

# Pioneer independent orbit determination

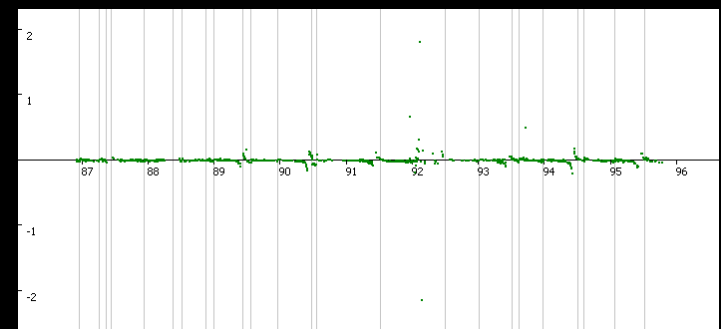
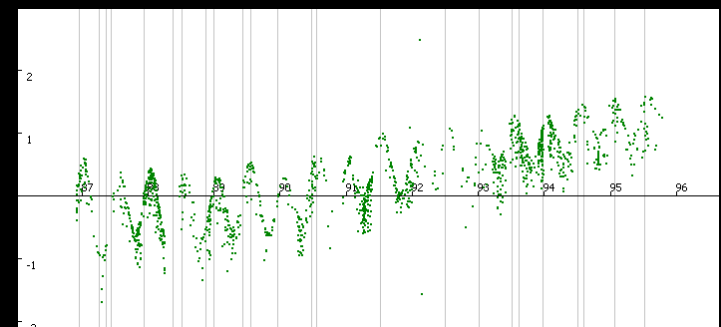
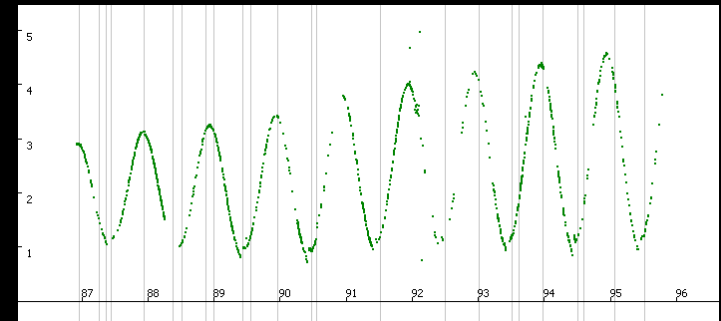
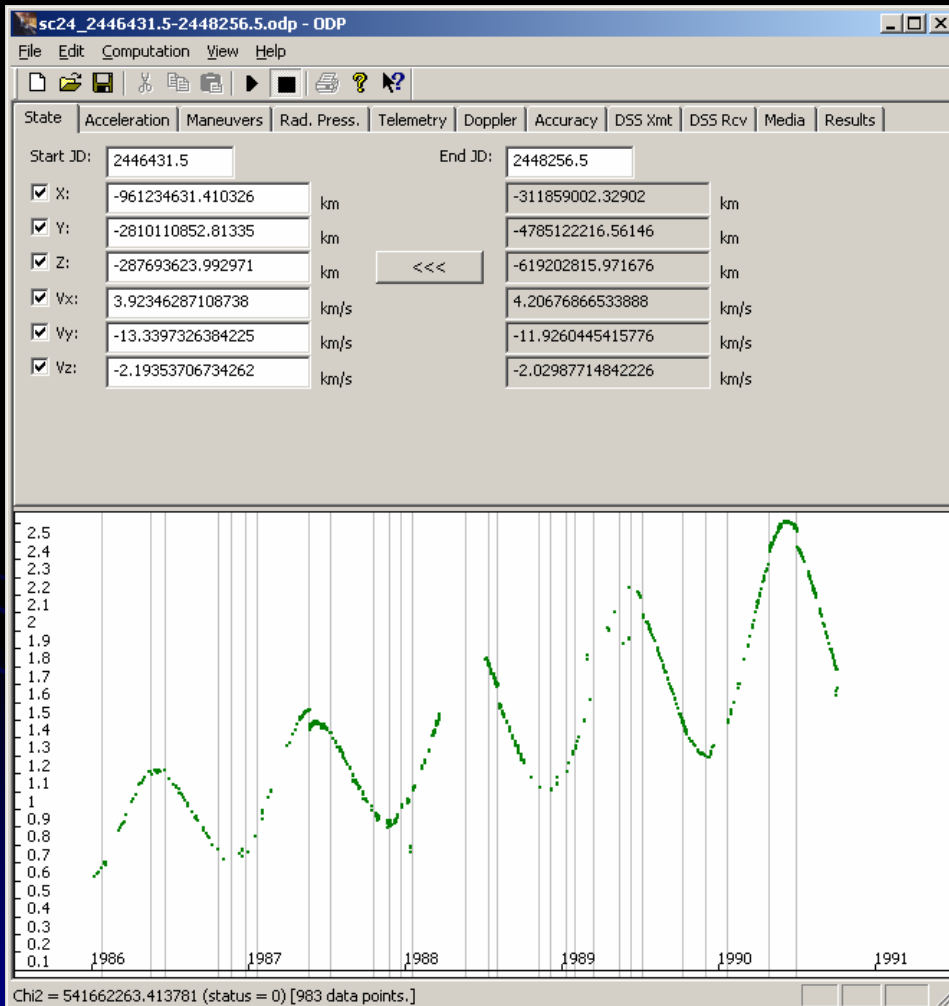
# Rationale

- Independent confirmation of anomalous acceleration
- More importantly: Ability to do precision orbit determination that incorporates telemetry-based estimation of on-board forces (or other theories!)

# Orbit estimation

- Simple orbit estimation program (C++, some FORTRAN libraries)
  - Modular construction: separate front-end, use of well-known libraries
  - Uses JPL SPICE library for solar system ephemerides
  - Takes into account current and historical DSN station positions
  - Uses MINPACK for nonlinear least squares estimation
  - Simple fast integrator (Stoermer's rule with predictor-corrector algorithm)
  - Takes into account relativistic Doppler and Shapiro time delay
  - Uses relativistic equations of motion for spacecraft position
  - Includes corrections due to oblateness of Jupiter and Saturn
  - Includes simple troposphere and solar corona model
  - Includes simple (Earth-pointing flat plate) solar pressure model
  - Uses custom library to read Doppler and ramp data from ODFs
  - Maneuvers estimated as changes in velocity along spin axis
  - Custom library reads telemetry for maneuvers, spin, and on-board generated heat
- JPL HORIZONS used for initial Pioneer 10/11 state estimates

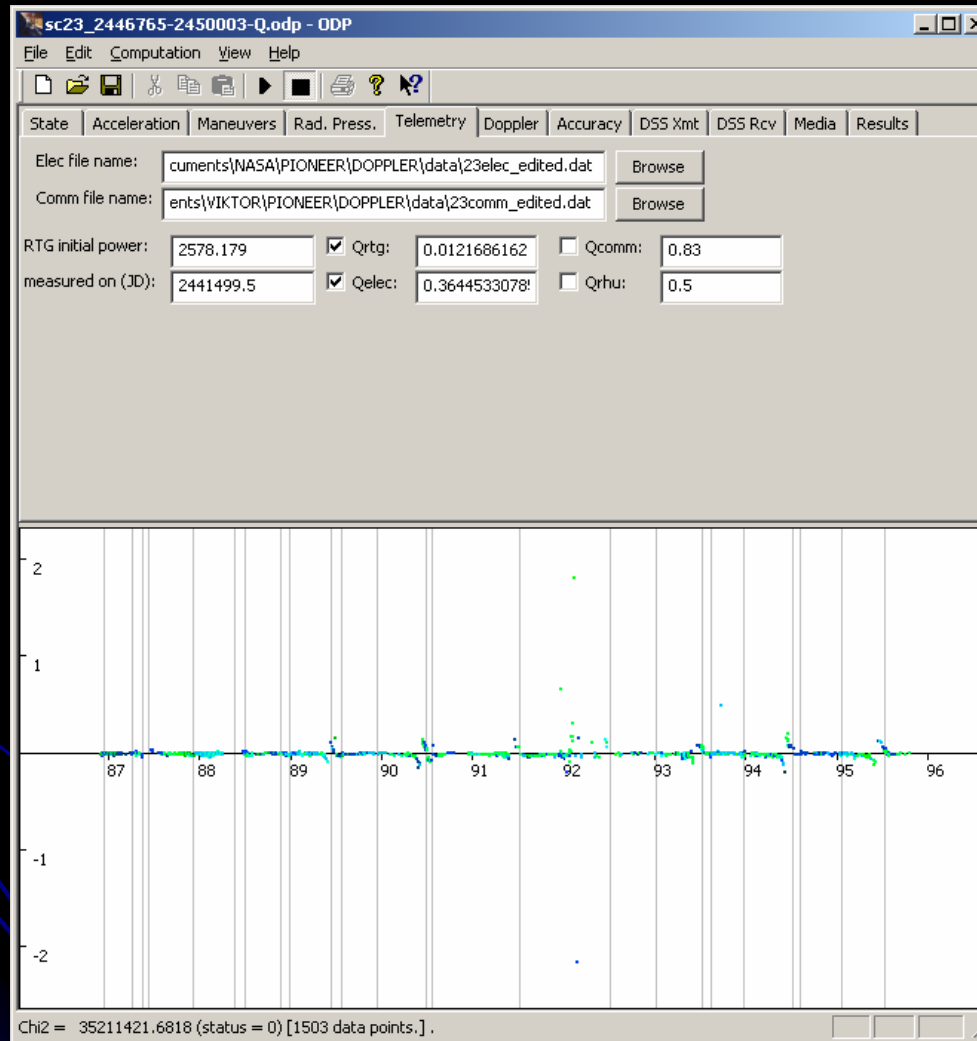
# Orbit estimation



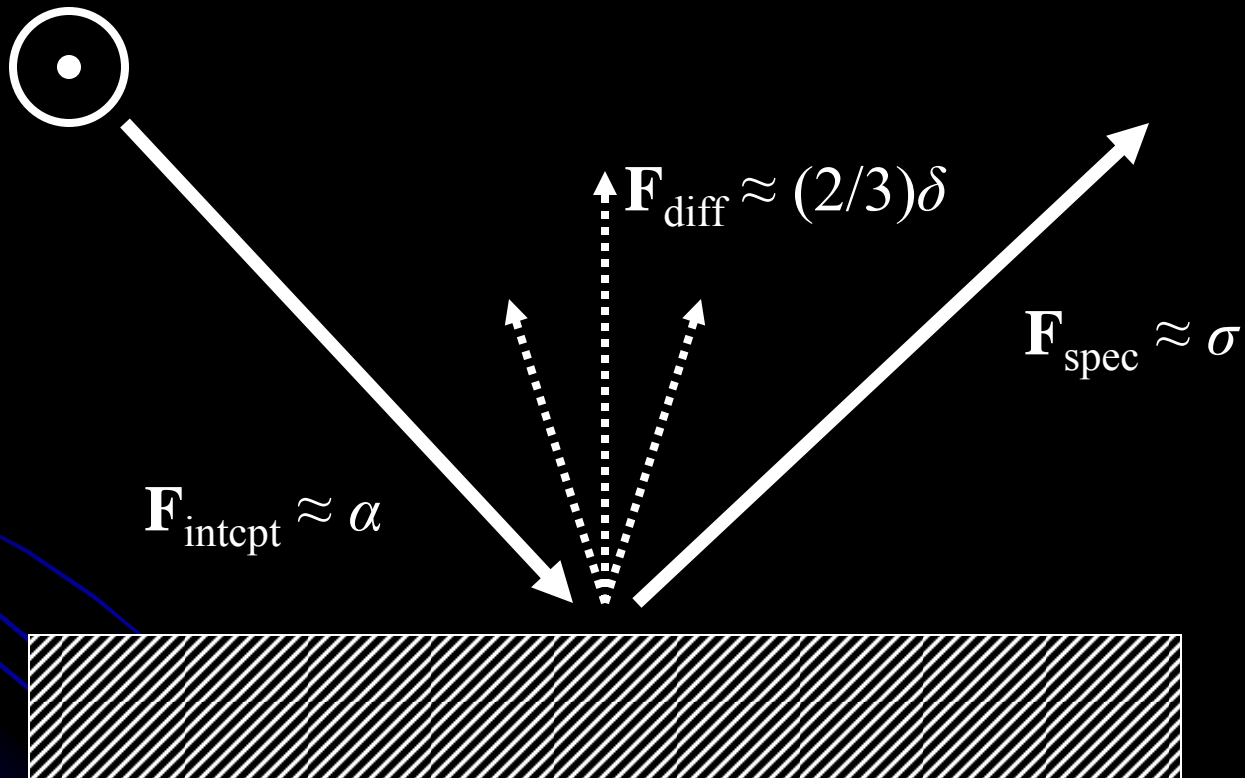
# Estimating on-board forces

- Basic concept: in addition to “solving for” orbital elements, also solve for linear factors ( $q_i$ ) of thermal acceleration. Equation of motion modified with thermal acceleration  $a = (1/mc)\Sigma q_i P_i$  where the  $P_i$  are known from telemetry.
- Method relies on no *a priori* assumptions about geometry, only about linearity and spin
- Most “correct” method from a modeling/statistical perspective

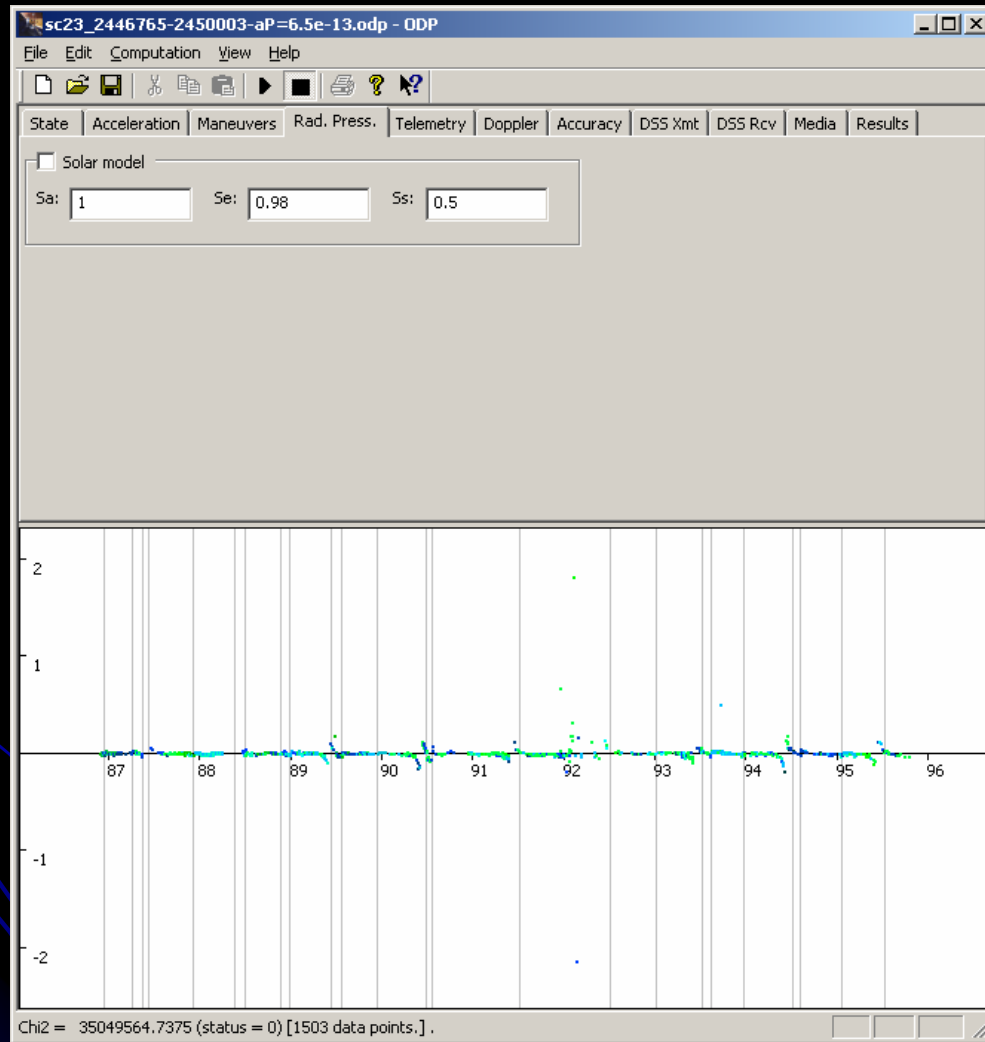
# Estimating on-board forces



# Solar model



# Solar model

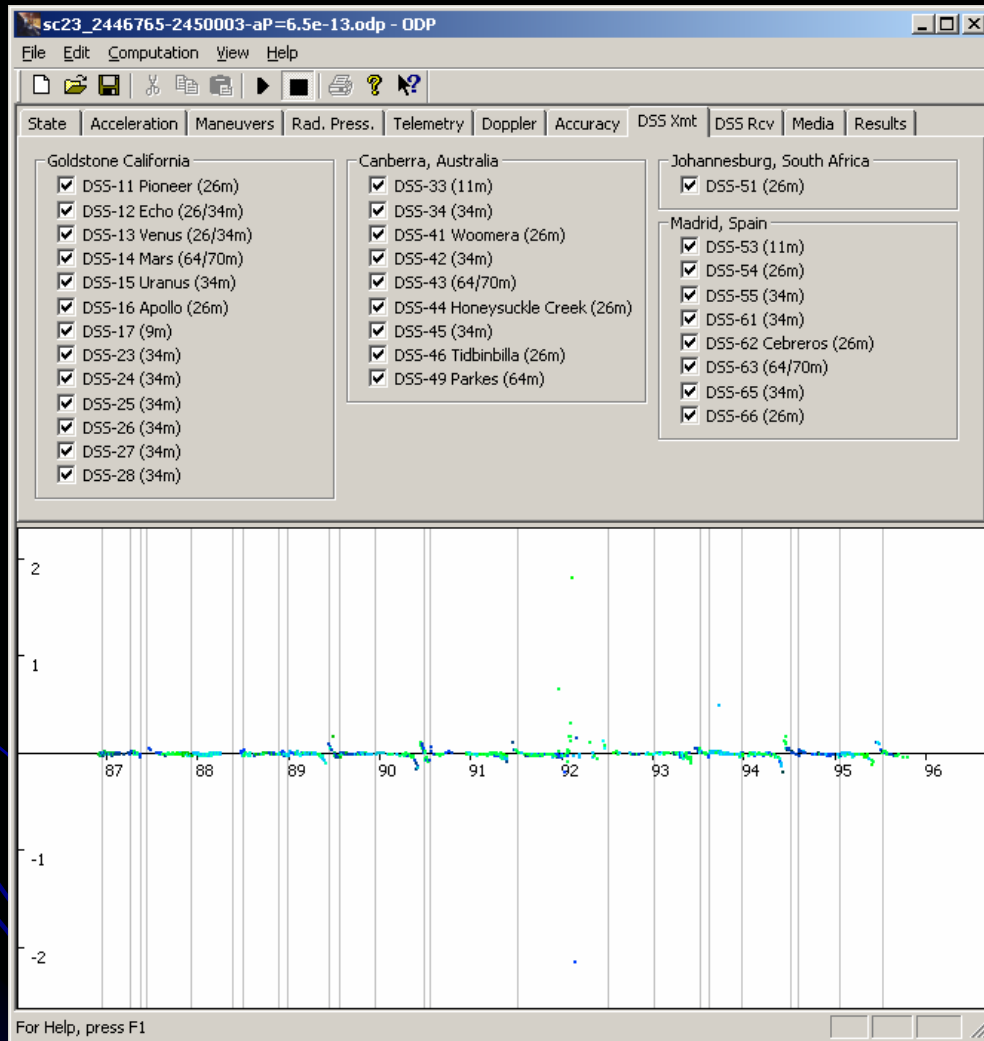




# DSN stations

- Pioneer data spans over 30 years
- During this period, many DSN stations have been
  - commissioned
  - decommissioned
  - moved
- Published data sets usually contain only current stations

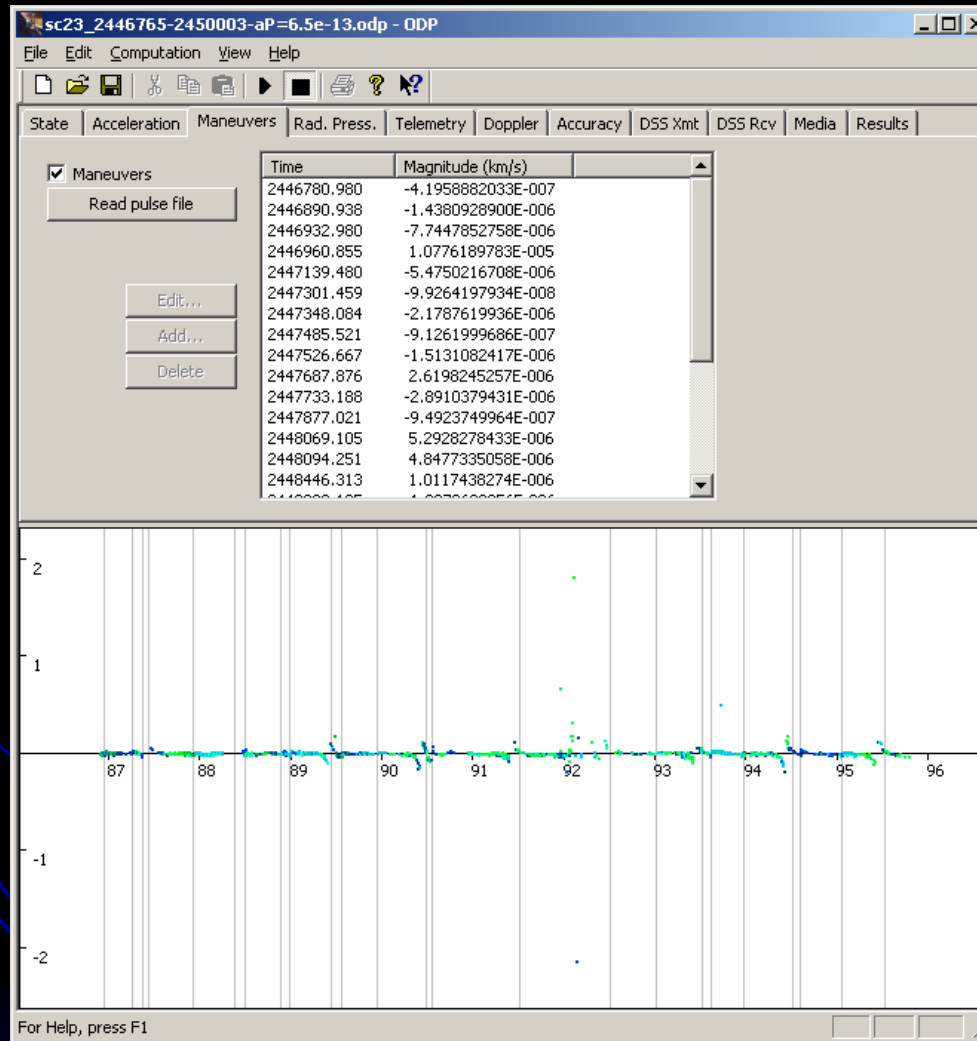
# DSN stations



# Maneuvers

- Attitude correction maneuvers are not supposed to change velocity; any velocity change is small and random (due to timing uncertainties, thruster alignment, etc.)
- Velocity changes perpendicular to the spin axis likely cancel each other as the spacecraft rotates
- Doppler observable not sensitive to such velocity changes anyway
- Instead of modeling each attitude correction as a three-dimensional maneuver, it is modeled as a velocity change along the Earth-spacecraft line, reducing the number of parameters to be estimated by a factor of three

# Maneuvers



# Data sets

- I had two data sets available:
- “JPL2002” certified data set used in the 2002 JPL study, containing:
  - Nearly 11 years of Pioneer 10 data (~9 years usable; last ~2 years use different encoding)
  - Nearly 5 years of Pioneer 11 data
- “Jacobson” uncertified data set containing:
  - ~75 days of data surrounding Pioneer 10 Jupiter encounter, ~67 days a year later
  - Nearly 2 years of data leading up to, and including, Pioneer 11 Saturn encounter

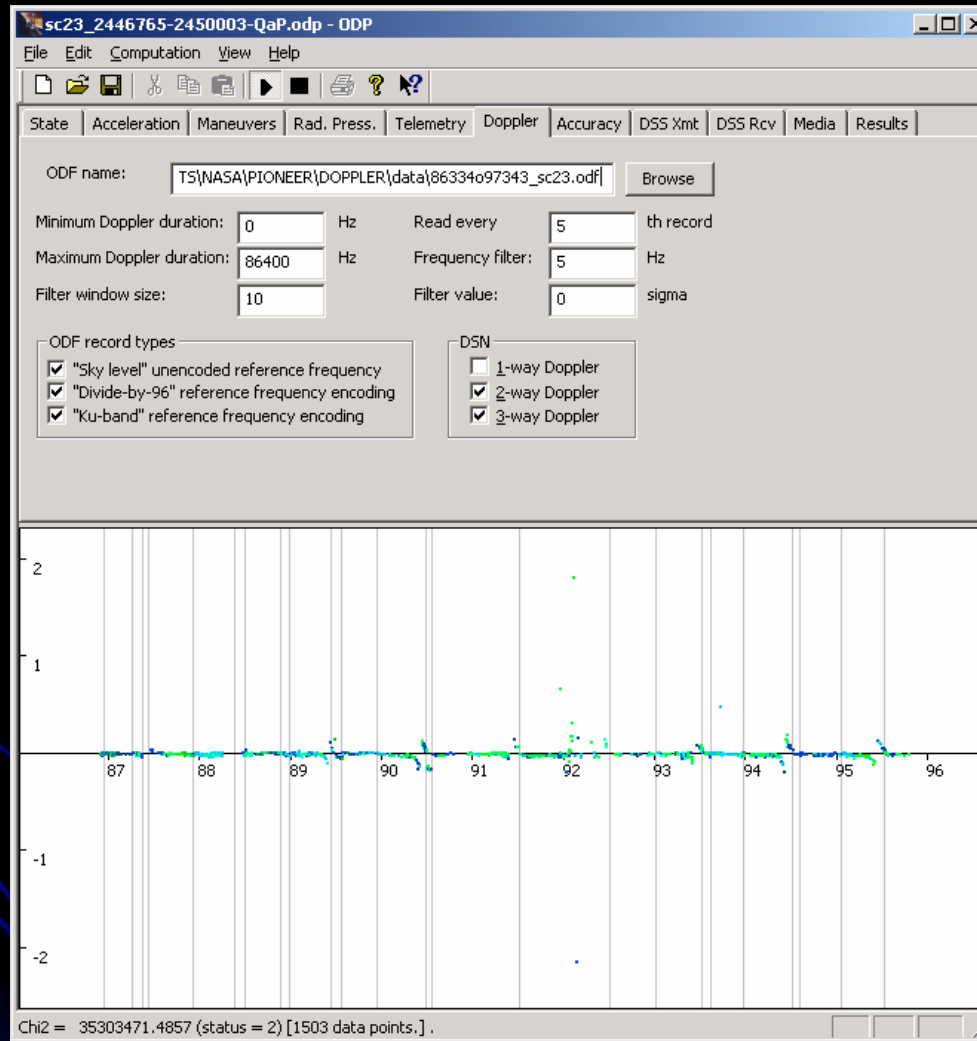
# Using data sets

- “Early” data not yet certified, but it can be used to
  - Validate orbit determination code
  - Verify/calibrate solar model
- “Late” (certified) data set can be used to
  - (Re)-determine the value of  $a_P$
  - Test the “thermal hypothesis”

# Some problems

- “Ku-band” data
- Sky frequency
- “Banding” when ramp data not used
- Block 5 data reference frequency
- Motion of Earth station during Doppler count interval
- Timekeeping

# Some problems





# Model validation

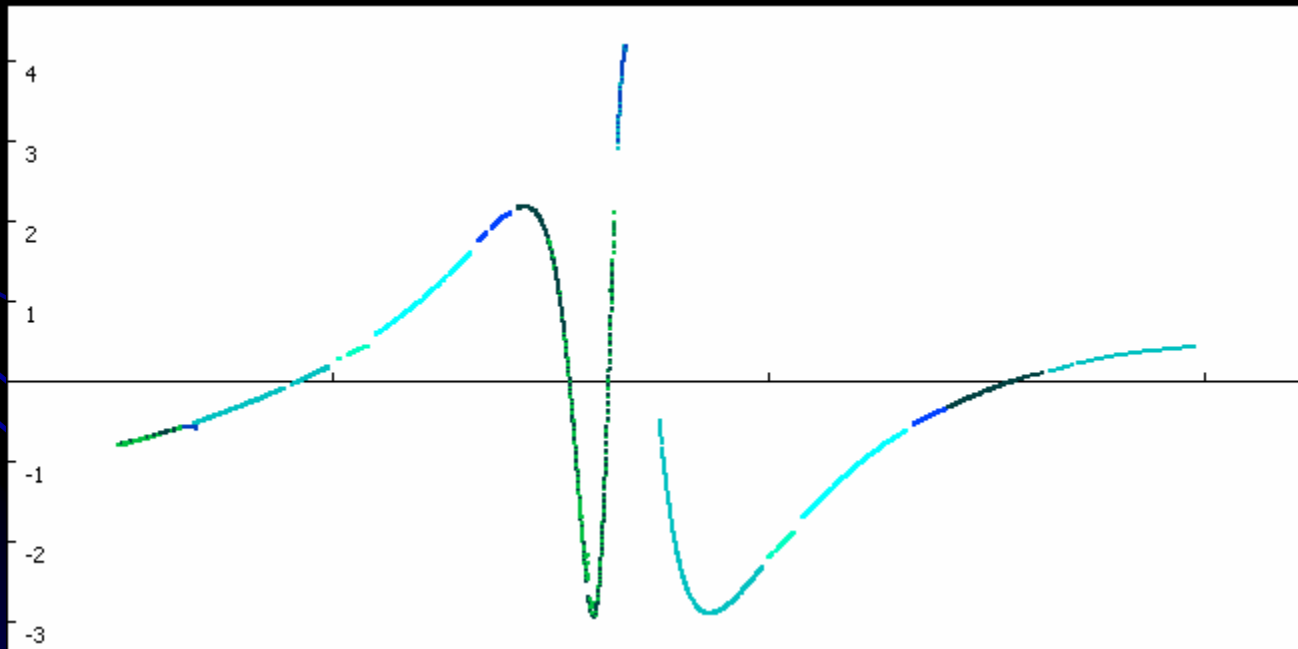
- Planetary encounter requires
  - Accurate orbit integration
  - Modeling of satellites
  - Relativistic equations of motion
  - Oblateness model
- Rapid change in spacecraft velocity stresses the model, allowing one to uncover subtle errors

# Model validation

- Planetary encounter not perfectly modeled; discrepancy due to
  - Not modeling the changing velocity of the spacecraft during Doppler interval
  - Not modeling planetary thermal radiation pressure
- Mismodeling is very similar for the two encounters

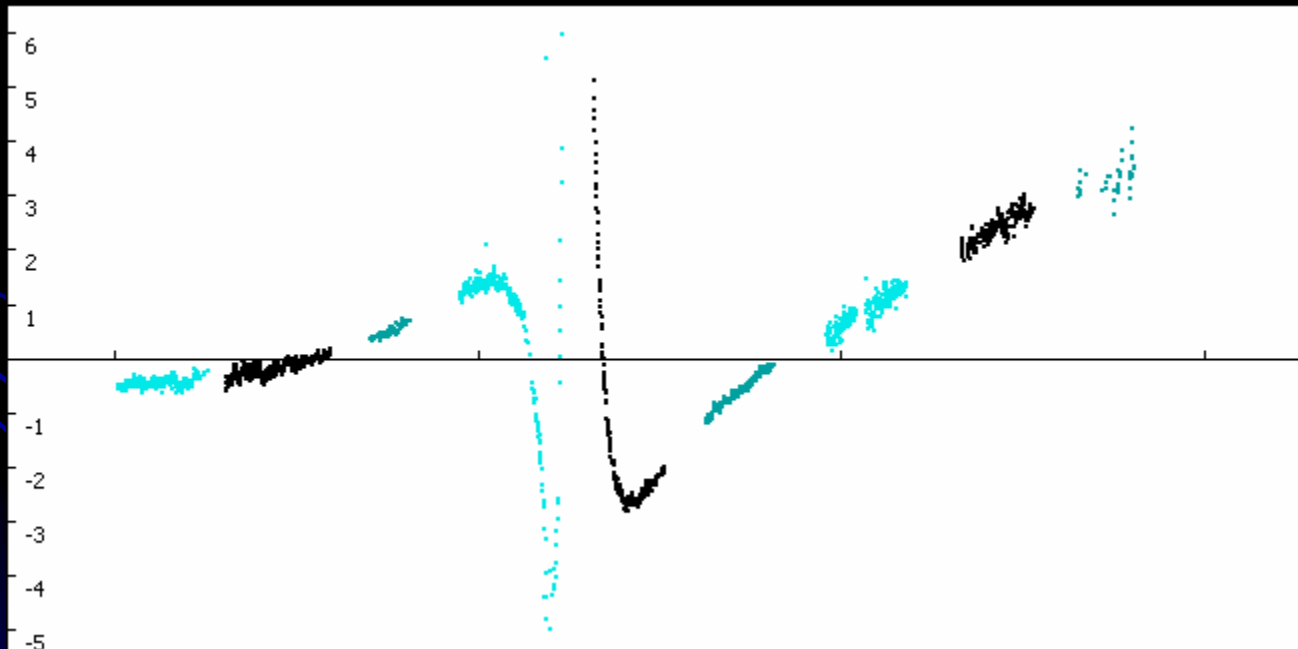
# Pioneer 10 Jupiter encounter mismodeling

- Mismodeling magnitude  $\sim 5$  Hz ( $\sim 30 - 40$  cm/s)
- Mismodeling limited to  $\sim 2$  day period surrounding closest approach



# Pioneer 11 Saturn encounter mismodeling

- Mismodeling magnitude  $\sim 5$  Hz ( $\sim 30 - 40$  cm/s)
- Mismodeling limited to  $\sim 2$  day period surrounding closest approach



# Solar model validation

- Simple 3-parameter flat plate model (for now) modeling absorption, diffuse and specular reflection

	$\alpha$	$\delta$	$\sigma$
Nominal	1.00	0.98*	0.50
Pioneer 10	1.00	0.77	0.56
Pioneer 11	0.95	1.34	0.43

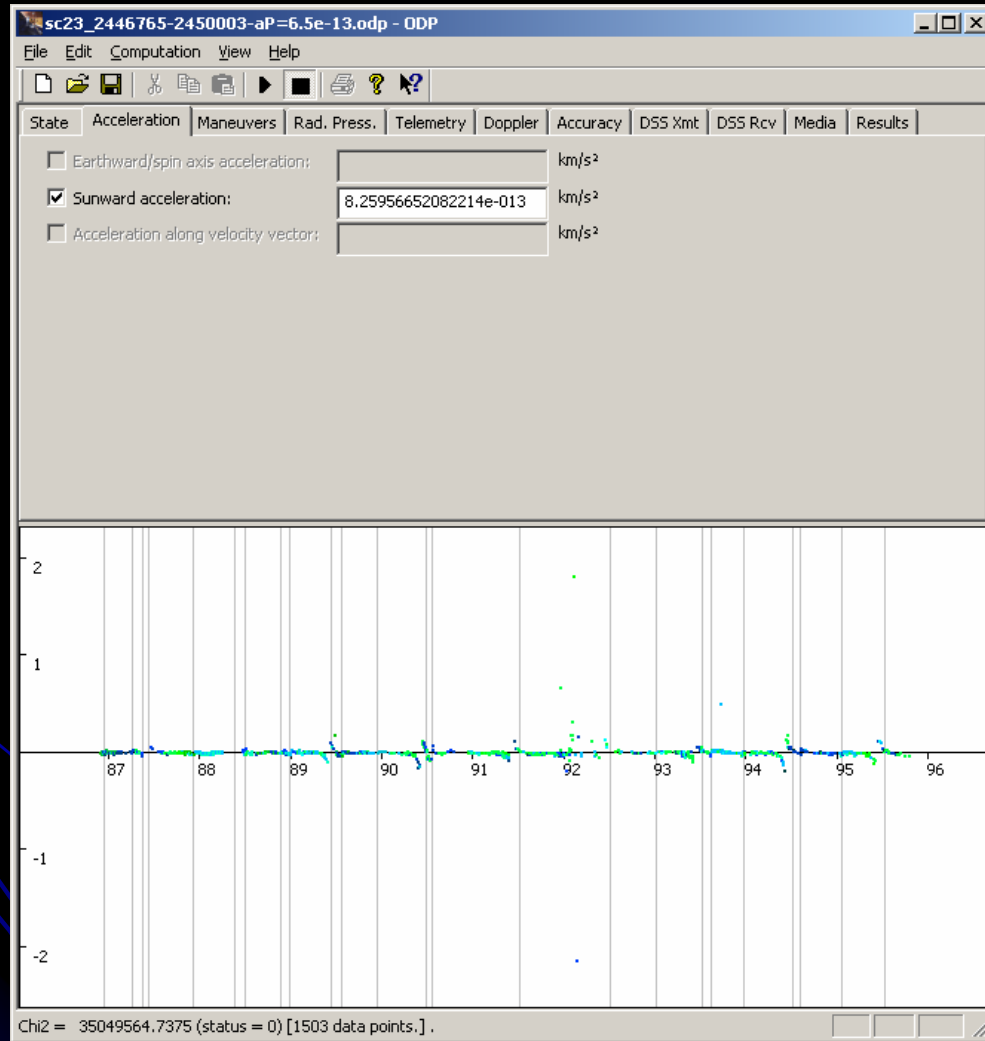
- Similarity between Pioneer 10 solution (<5 AU) and nominal values increases confidence that solar model is valid and orbit determination code works reliably

\*Published reflectance of 0.79 is assumed to be “net” reflectance, from this and emittance figures, an effective reflectance of 0.96 is calculated. Results presented here used the figure of 0.98 (a result of a prior calculation); the difference between the two is negligible.

# Solving for $a_p$

- Solution appears highly dependent on initial value (least squares optimization finds local minima near initial value; acceleration absorbed by maneuvers)
- The goodness of fit ( $\chi^2$ ) is similar in all cases
- Therefore, I cannot say that a particular value of  $a_p$  is “the” solution, only that it is consistent with the data

# Solving for $a_p$



# Solving for $a_p$

- All values in units of  $10^{-13}$  km/s<sup>2</sup>:

Data set	Initial value:	0	10	6.5
"2002"	Pioneer 10	4.77	8.06	8.26
	Pioneer 11	-0.15	7.38	5.84
"Jacobson"	Pioneer 11	-0.29	9.55	5.99

- **Conclusions so far (this is work in progress!): much depends on the choice of initial values and the least squares algorithm**
- **Same near constant acceleration consistent with all studied data sets**



# Solving for $a_p$ and $q_i$

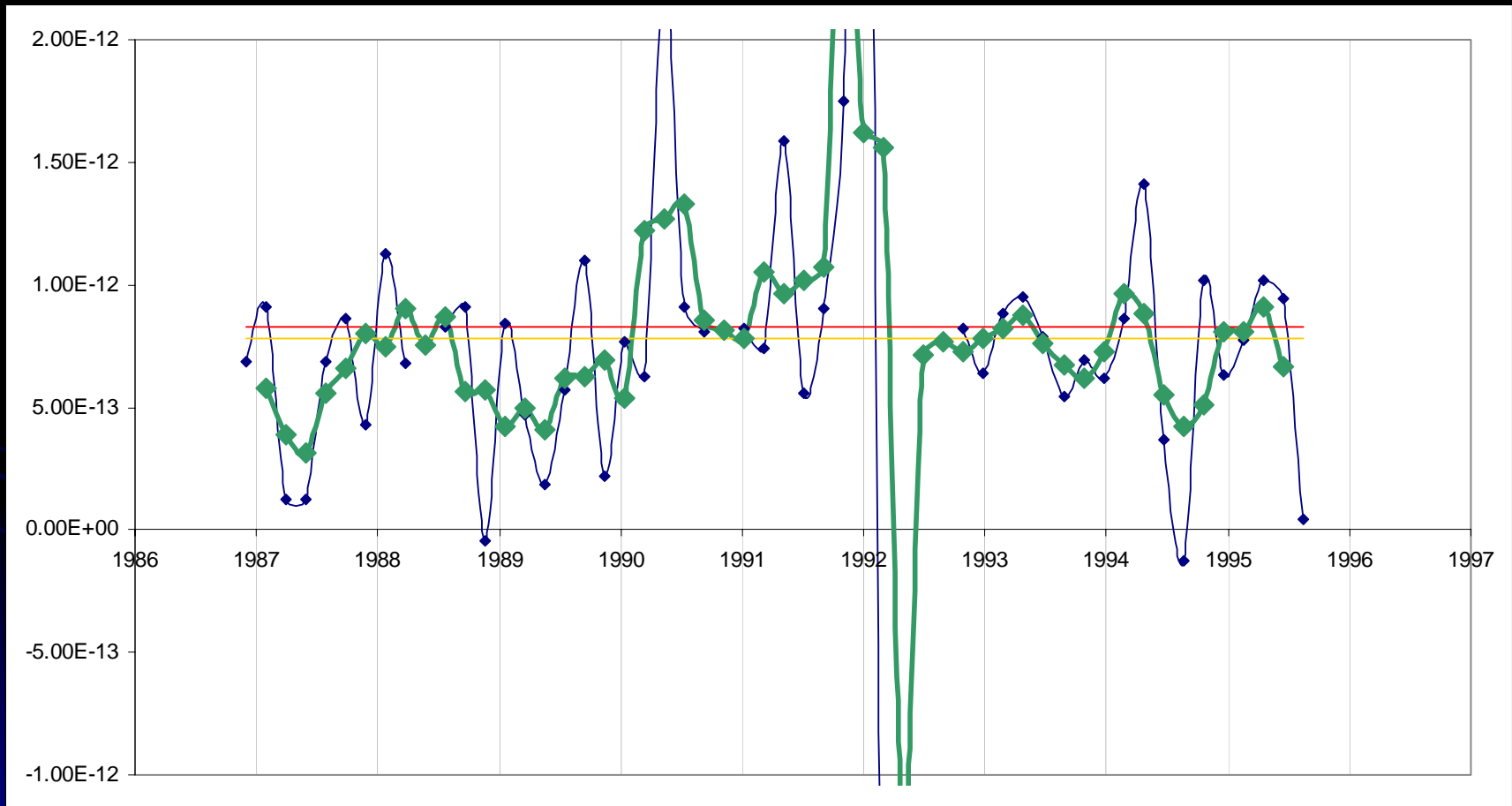
- Simultaneously solving for constant acceleration and thermal coefficients
- Initial thermal values are nominal:  
 $q_{\text{RTG}} = 0.01$ ,  $q_{\text{elec}} = 0.35$
- Two different initial accelerations used:

Initial $a_p$ ( $10^{-13}$ km/s <sup>2</sup> )	$a_p$	$q_{\text{RTG}}$	$q_{\text{elec}}$
0	0.15	0.012	0.36
4	2.05	0.010	0.36

# Stochastic acceleration

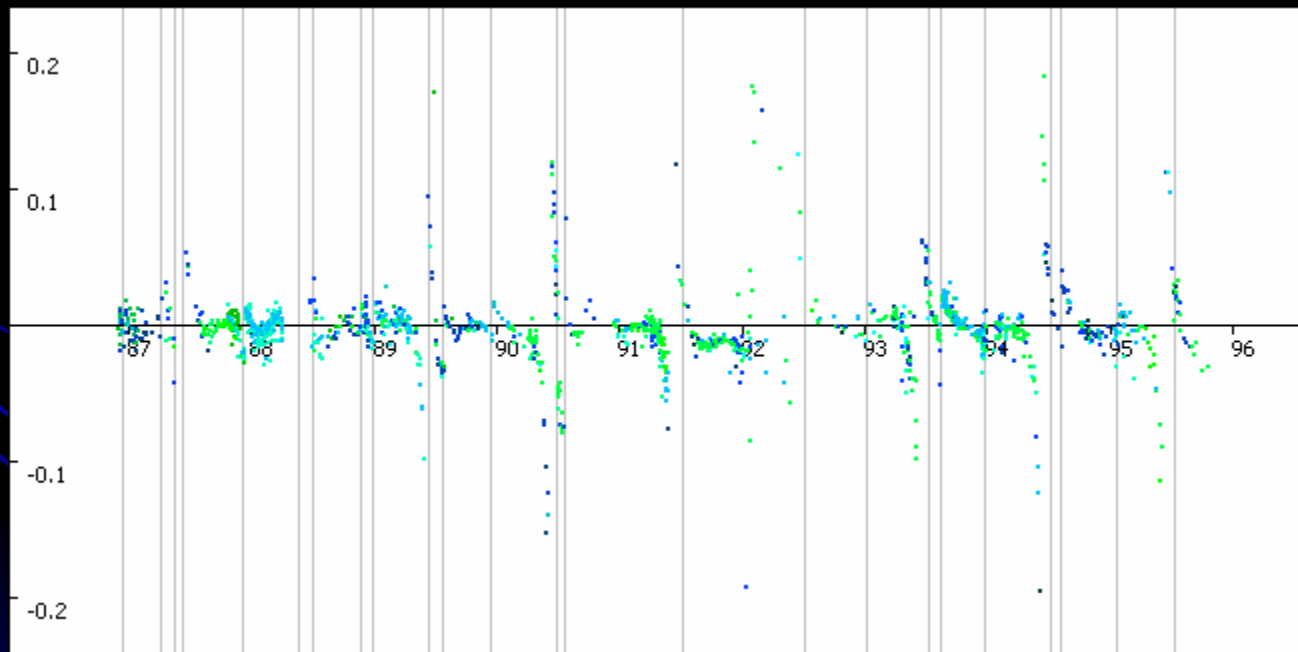
- After obtaining initial solution:
  - Divide data into 60-day (120-day) batches
  - Propagate solution through batches, solving for only
    - $a_p$
    - Maneuvers
- Used 9-year Pioneer 10 data set

# Stochastic acceleration



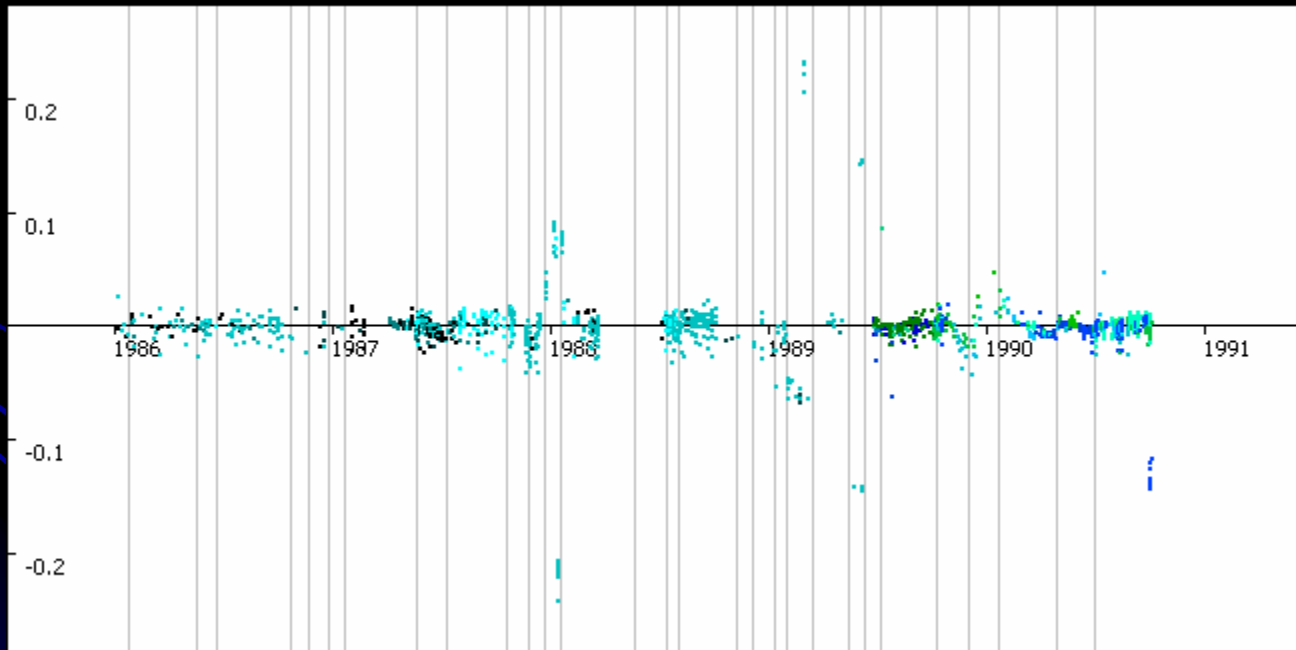
# Stochastic acceleration

- Some annual periodicity is also visible in least squares solution (error in Shapiro code?)



# Stochastic acceleration

- Same signature not apparent in Pioneer 11 data set



# Plans (February 2007)

- Develop proper three-dimensional radiation pressure model using raytracing code
- Improve orbit determination code:
  - Improve flyby modeling accuracy
  - Solar and planetary radiation model
  - Ionospheric propagation delay
- Incorporate means to estimate acceleration (and jerk term!) in other directions (Earth-pointing, and along the direction of motion)
- Improve attitude and maneuver modeling
- Improve pointing information using AGC or SNR
- Improve data weighing, filtering, etc. capabilities
- FIND AND FIX BUGS!
- Utilize new Doppler data when available

# Thank You!